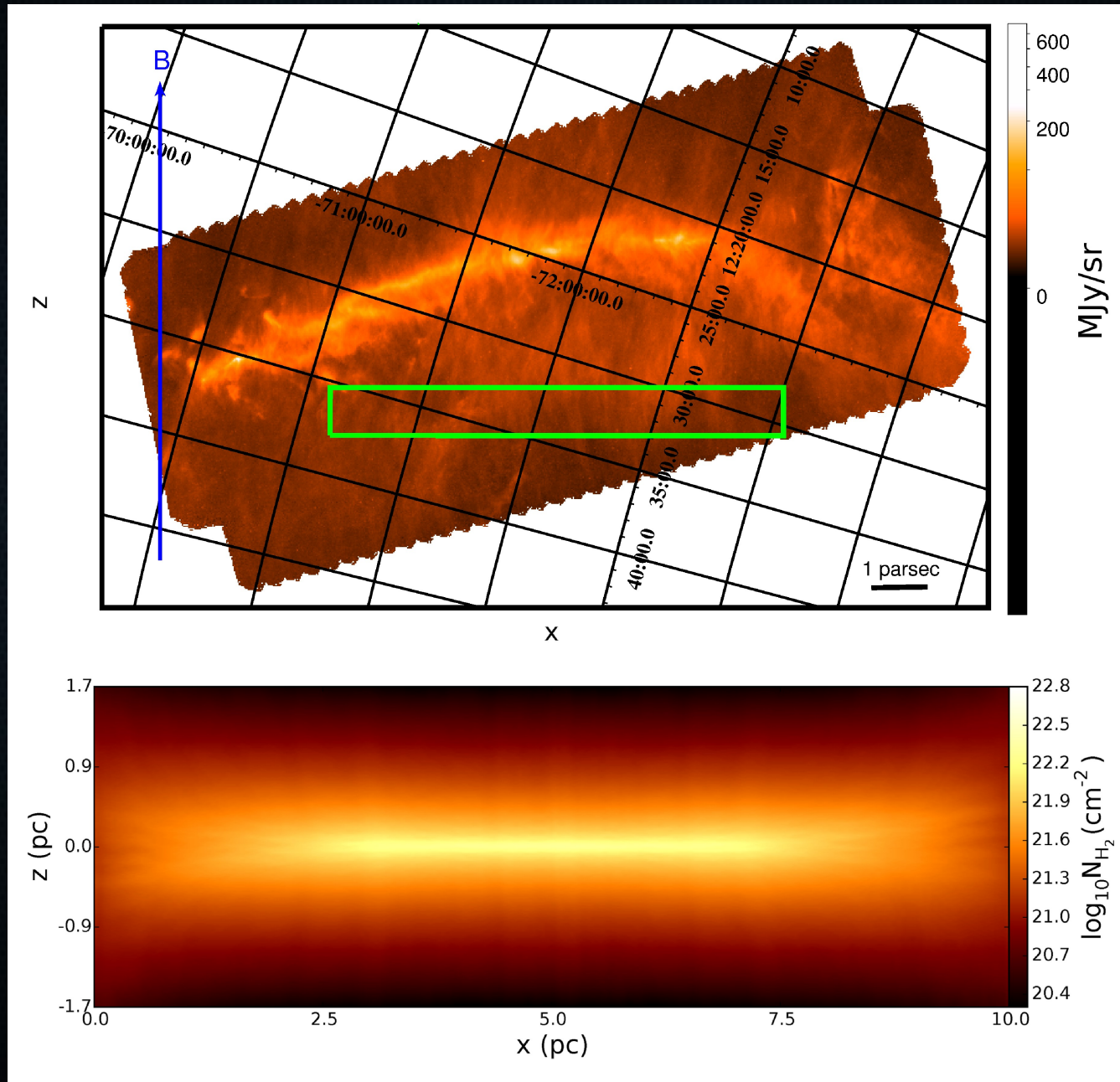
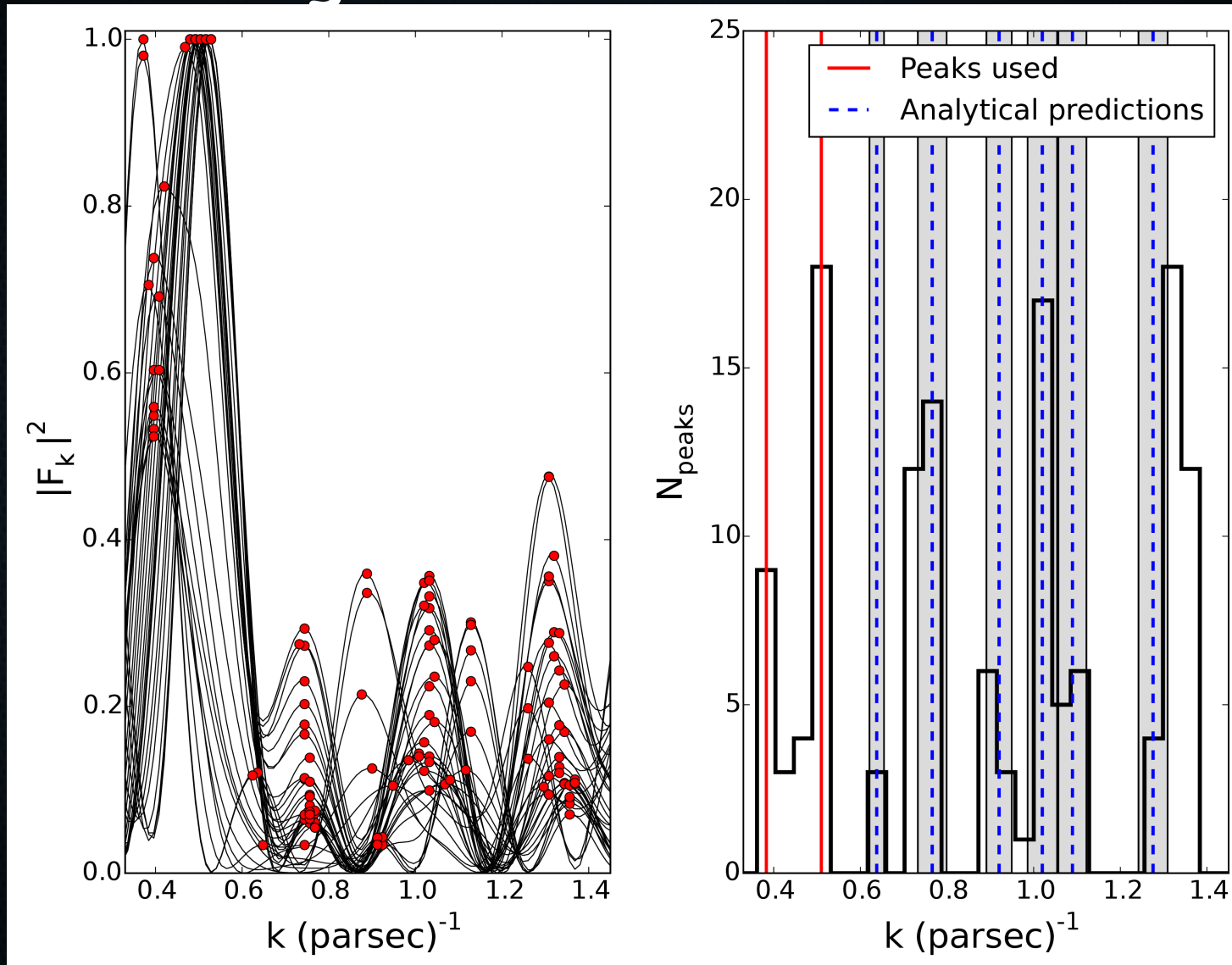


Magnetic seismology of interstellar clouds: unveiling the hidden dimension

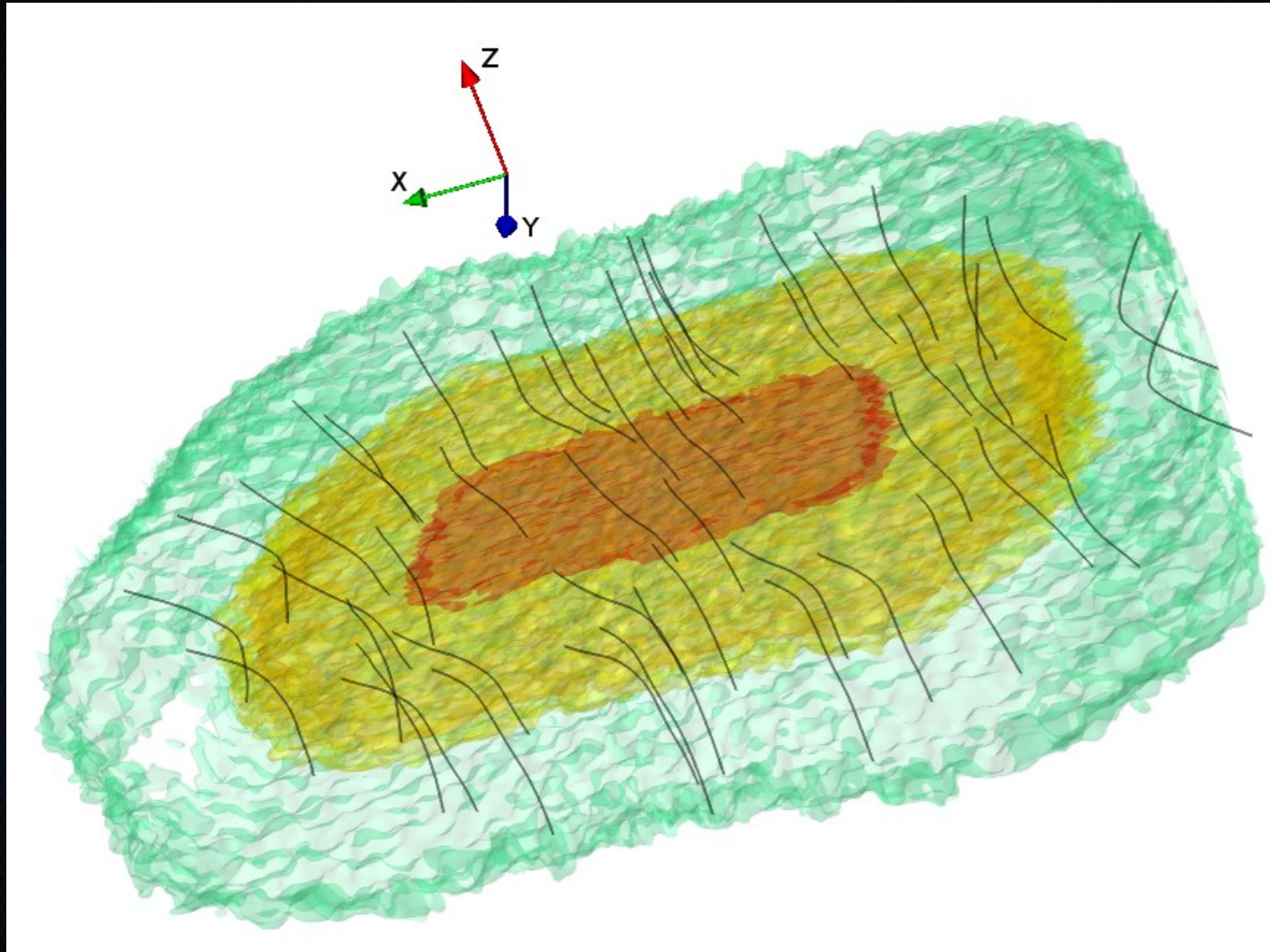


Magnetic seismology of interstellar clouds: unveiling the hidden dimension



$$k_{mn} = \sqrt{\left(\frac{\pi n}{L_x}\right)^2 + \left(\frac{\pi m}{L_y}\right)^2}$$

Magnetic seismology of interstellar clouds: unveiling the hidden dimension



Hidden - LOS dimension = 6.4 pc => Musca
sheet viewed edge-on

Next step

Calculate the magnetic field strength from the dispersion relation of FMS waves

$$v_{\phi}^2 = (\omega/k)^2 \Rightarrow v_A^2 + c_s^2 = (2\pi/kT)^2 \Rightarrow \\ B^2 = 16\pi^3 \rho / (kT)^2 - c_s^2 4\pi\rho$$

For $B \nearrow$ and $\rho \searrow$ $T \sim 60$ years at the resolution limit (optical)

Early observations (photographic plates) ~ 60 years old!

Best way to get velocity dispersion from double peaked line profiles?