Astrophysical Gas Dynamics

NEXT:

- *Hydrodynamical Equations (recap → notes)*
- *The Equation of State (→ notes)*
- *Derivation of the Energy Equation of Hydrodynamics (→ notes)*

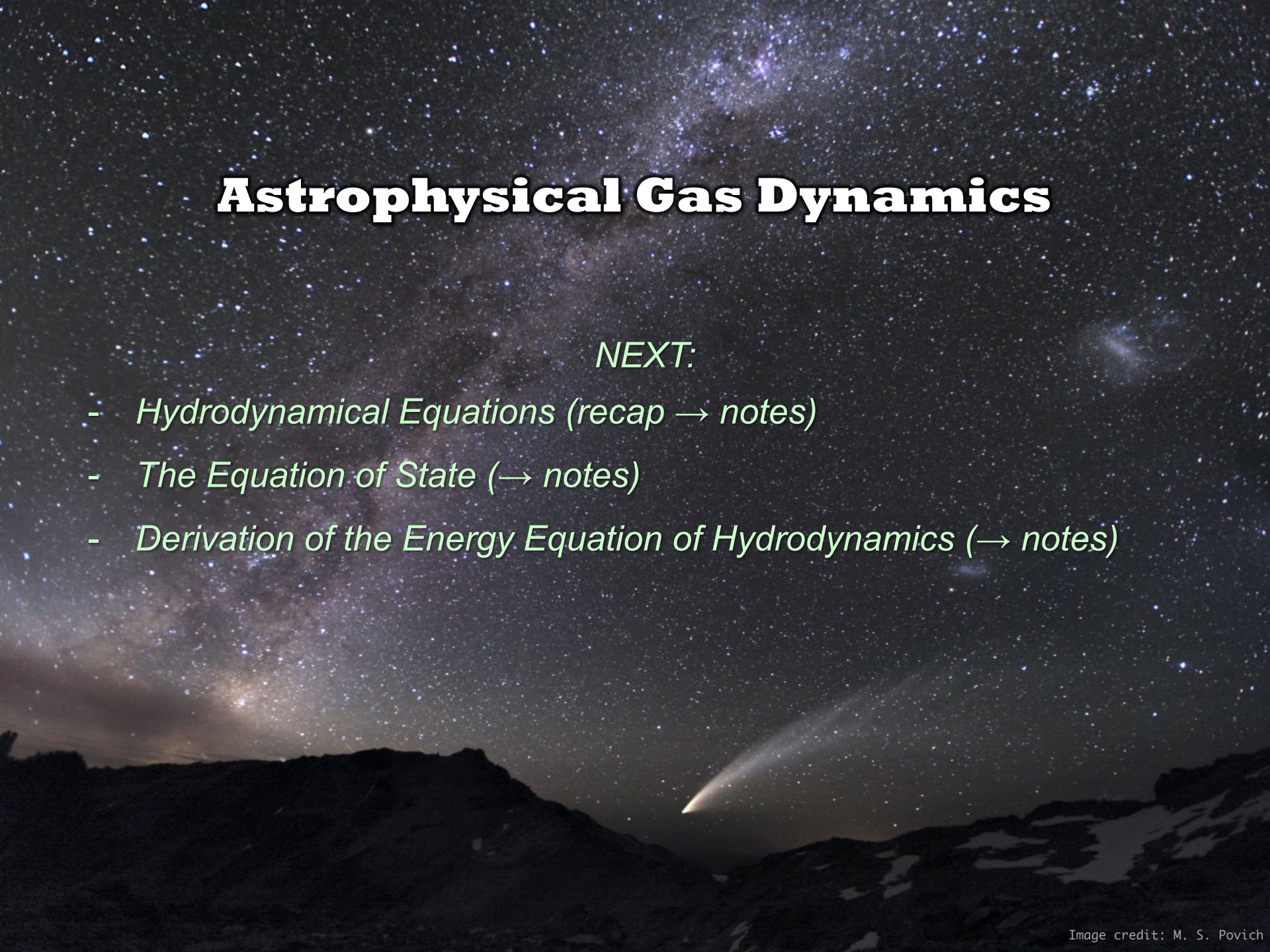
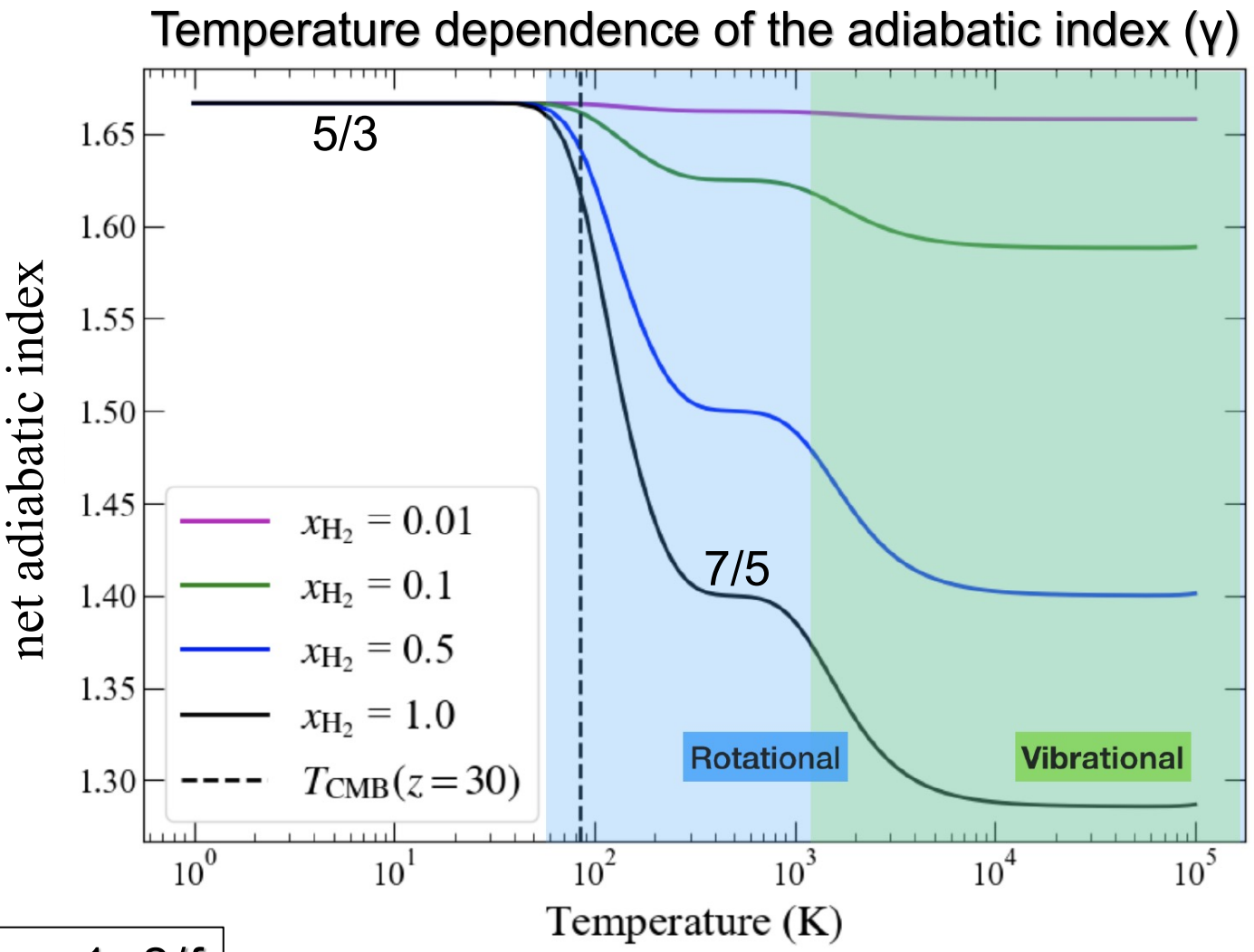


Image credit: M. S. Povich

Adiabatic Index



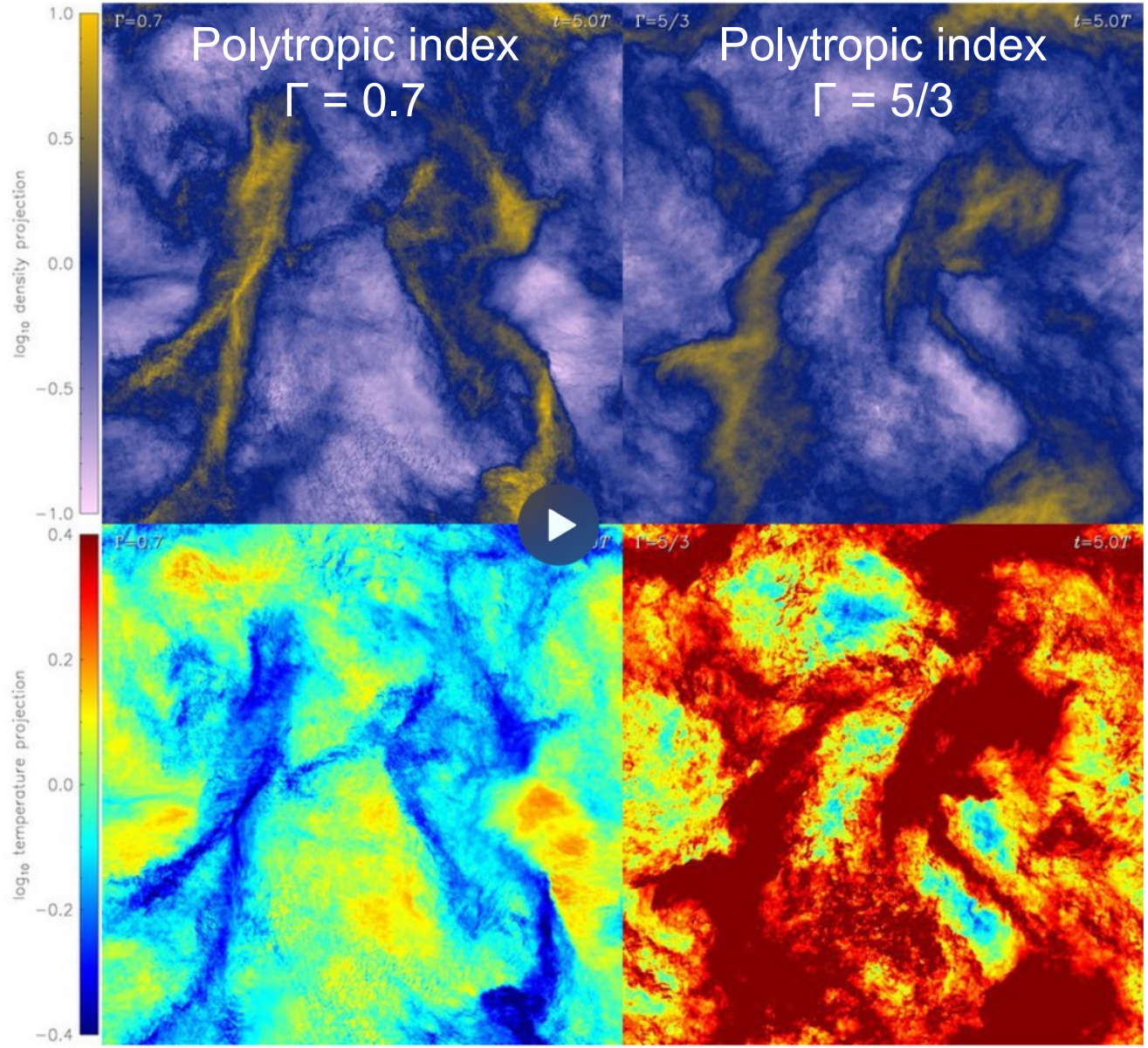
$\gamma = c_P/c_V = 1+2/f$

Number of degrees of freedom (f) depends on excitation of rotational and vibrational states (partition functions: T-dependence).

Sharda et al. (2019)

Equation of State – Polytropic EOS

$$P_{\rm th} = K \rho^{\Gamma}$$



Density

Temperature

Movies available: http://www.mso.anu.edu.au/~chfeder/pubs/polytropic/polytropic.html

Federrath & Banerjee (2015)

Energy equation with heating and cooling

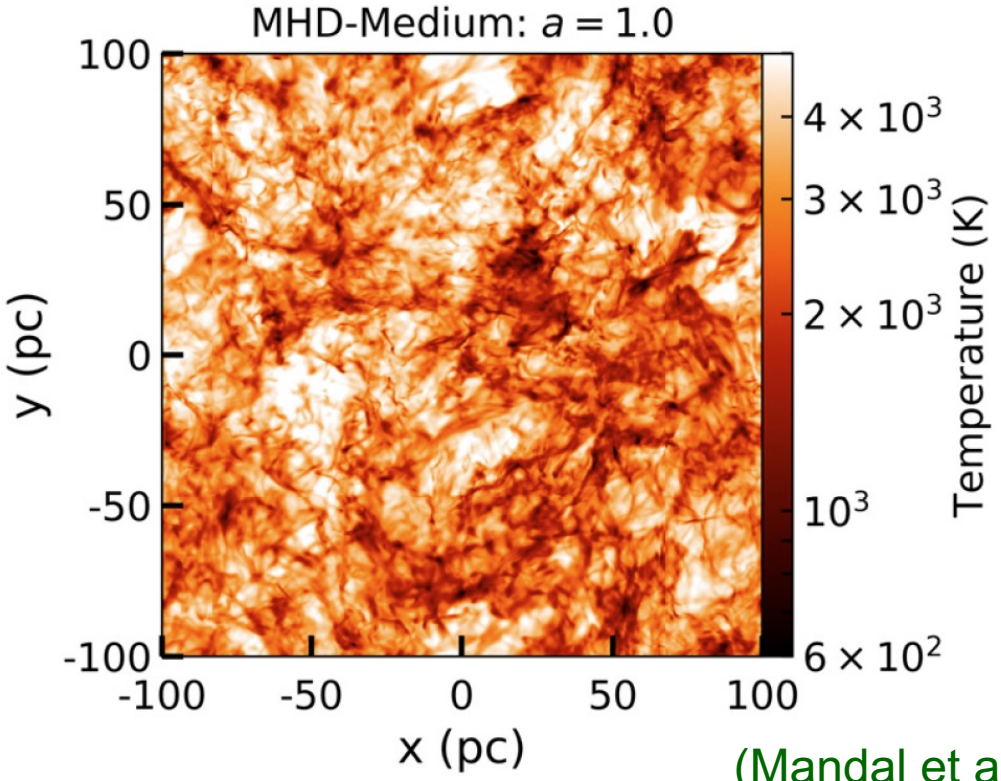
$$\frac{\partial}{\partial t} e_{\rm tot} + \nabla \cdot \left[(e_{\rm tot} + P_{\rm tot}) \boldsymbol{v} - \frac{1}{4\pi} (\boldsymbol{B} \cdot \boldsymbol{v}) \boldsymbol{B} \right] = \frac{1}{\rho} \left[\frac{\rho}{\mu m_{\rm H}} \Gamma - \left(\frac{\rho}{\mu m_{\rm H}} \right)^2 \Lambda(T) \right]$$

Heating:

$$\Gamma = 2 \times 10^{-26} \, {\rm erg \, s^{-1}}$$

Cooling:

$$\frac{\Lambda(T)}{\Gamma} = 10^7 \exp\left(\frac{-1.184 \times 10^5}{T + 1000} \right) + 1.4 \times 10^{-2} \sqrt{T} \exp\left(\frac{-92}{T} \right) {\rm cm^3}$$



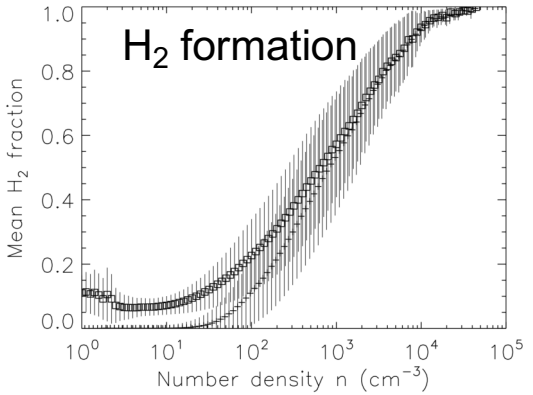
- Photoelectric heating from small grains and polycyclic aromatic hydrocarbons (PAHs)
- Heating and ionization from cosmic rays and X-rays
- H_2 formation and destruction
- Atomic and molecular line cooling

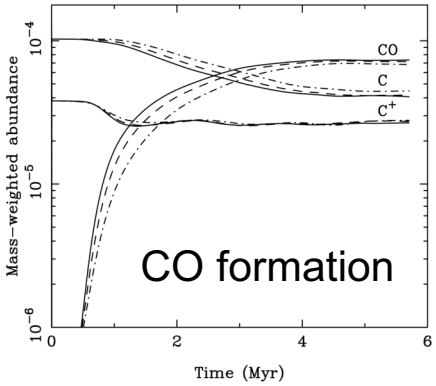
(Sutherland & Dopita 1993; Koyama & Inutsuka 2002; Vazquez-Semadeni et al. 2007)

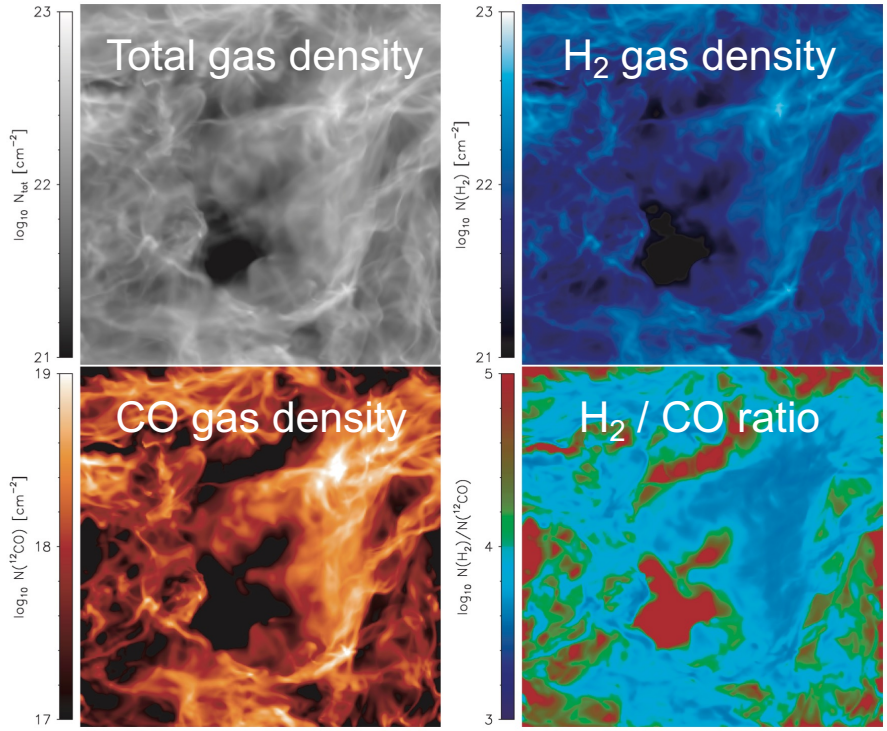
(Mandal et al. 2020)

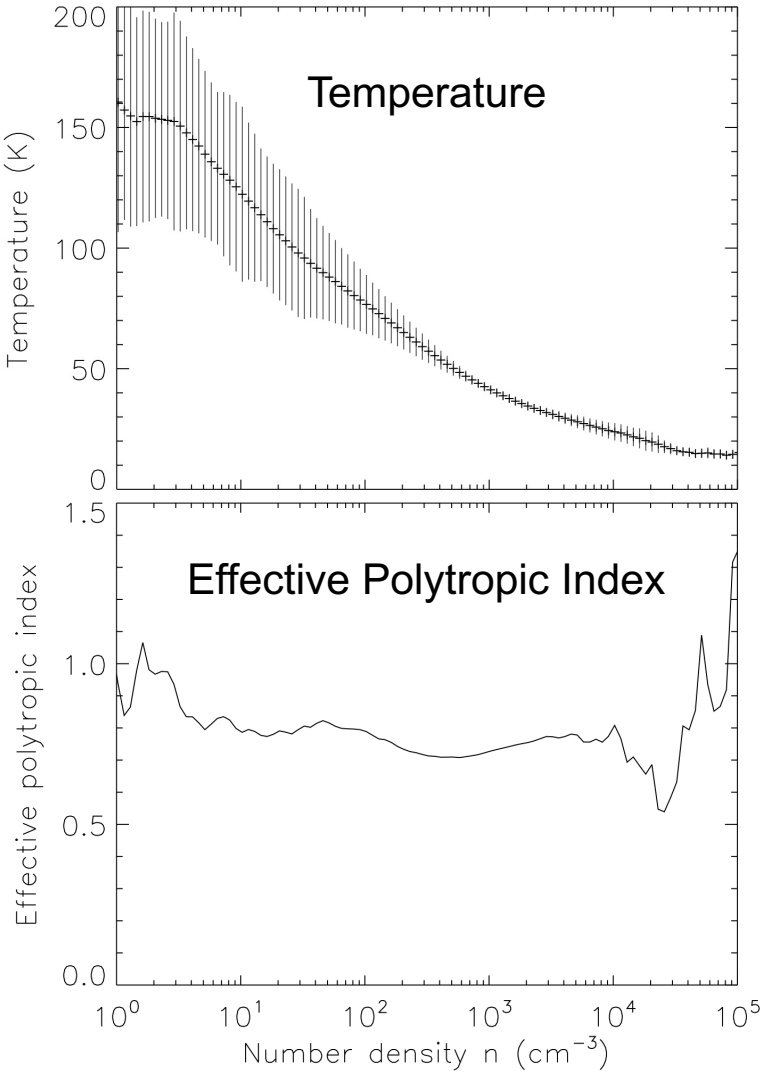
Equation of State – Chemistry / Heating / Cooling

Chemistry / Heating / Cooling: (Glover et al. 2007, 2010)



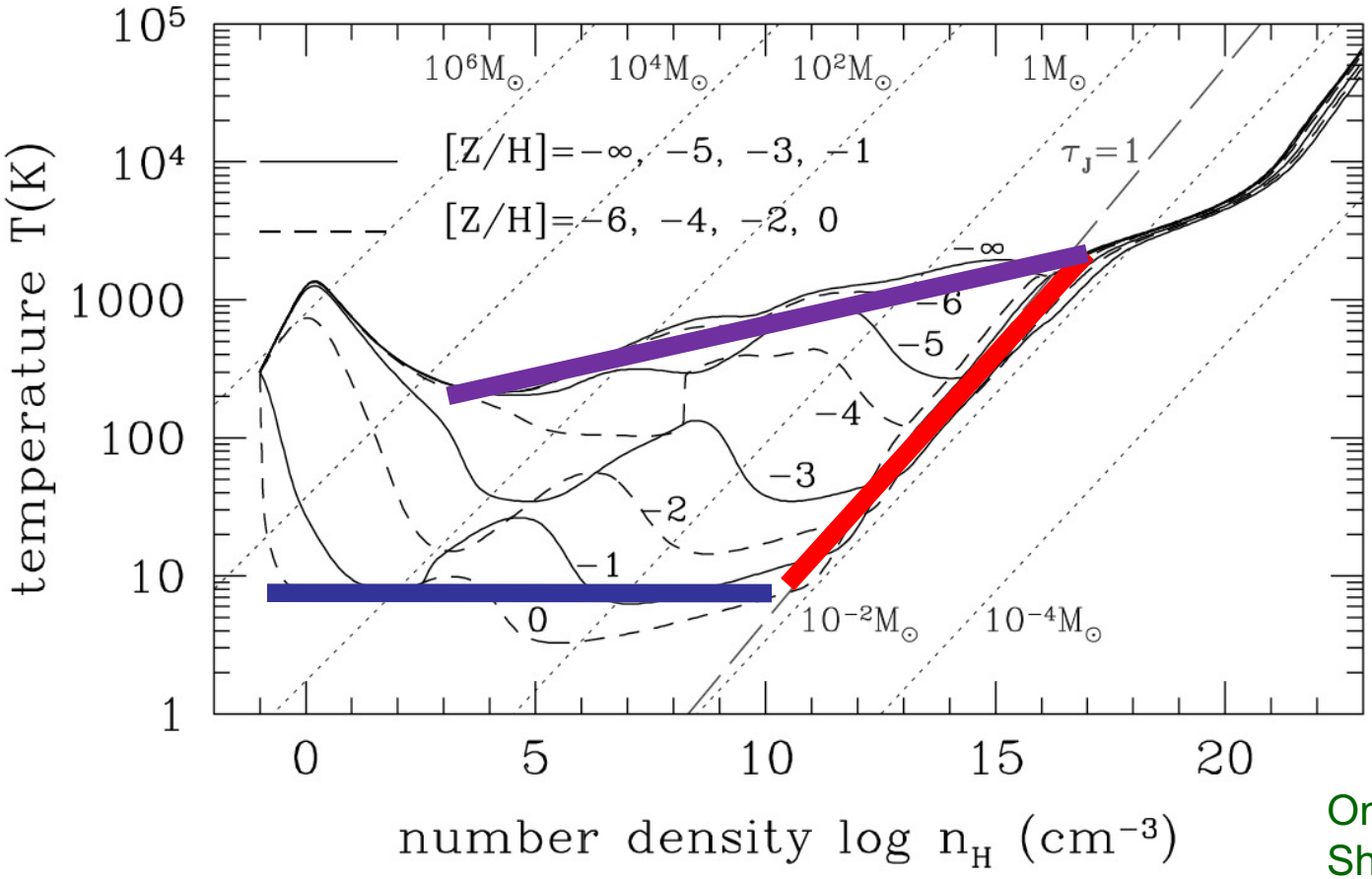






Equation of State – Chemistry / Heating / Cooling

Chemistry / Heating / Cooling: (Glover+07,10, Micic+12, Clark+12)



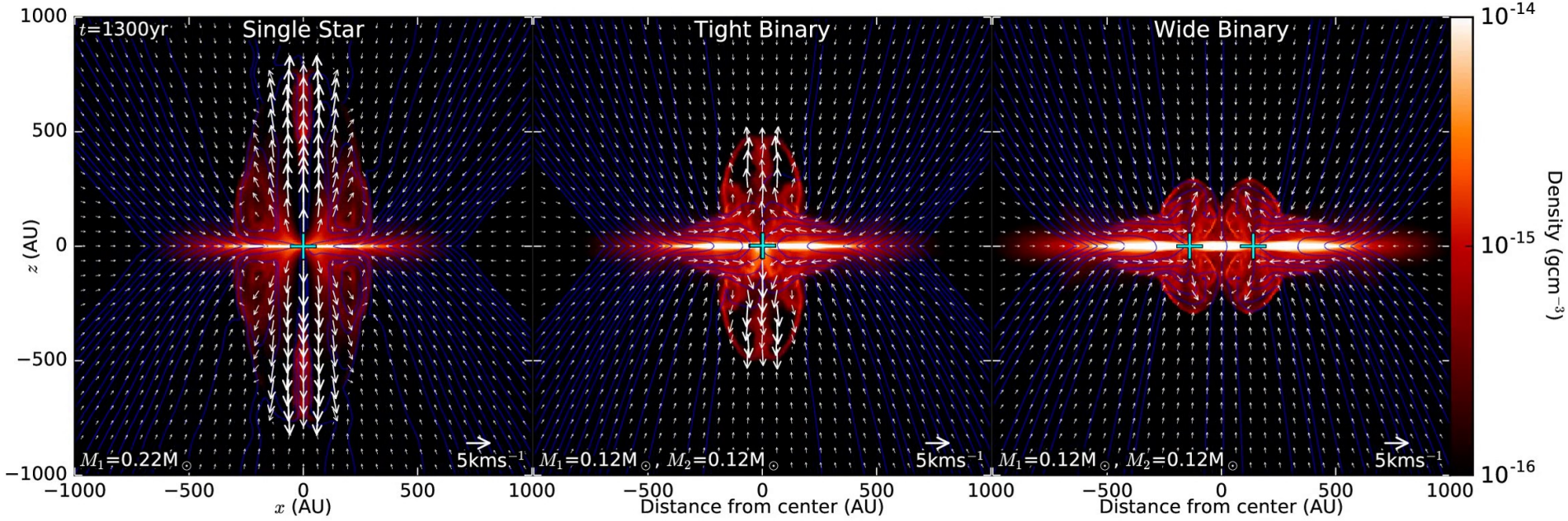
Omukai et al. (2005); Sharda et al. (2020)

Molecule formation in high-density gas: $t_{form} \sim 1/n$ Micic et al. (2012), Hollenbach et al. (1971)

Equation of State – Polytropic EOS for Star Formation

$$P_{\rm th} = K\rho^{\Gamma} \quad \text{with} \quad \Gamma = \begin{cases} 1 & \text{for} \quad \rho \le \rho_1 \equiv 2.50\times10^{-16}\,{\rm g\,cm^{-3}}, \\ 1.1 & \text{for} \quad \rho_1 < \rho \le \rho_2 \equiv 3.84\times10^{-13}\,{\rm g\,cm^{-3}}, \\ 1.4 & \text{for} \quad \rho_2 < \rho \le \rho_3 \equiv 3.84\times10^{-8}\,{\rm g\,cm^{-3}}, \\ 1.1 & \text{for} \quad \rho_3 < \rho \le \rho_4 \equiv 3.84\times10^{-3}\,{\rm g\,cm^{-3}}, \\ 5/3 & \text{for} \quad \rho > \rho_4\,. \end{cases}$$

Movies available:
https://www.mso.anu.edu.au/~chfeder/pubs/binary_jets/binary_jets.html



Kuruwita et al. (2017); Gerrard et al. (2018)