Astrophysical Gas Dynamics Basic Equations of Hydrodynamics - Introduce gas/fluid variables: density S, velocity V, energy e, pressue P, temperature T ... ave related by the Hydros egs. and the Equation of State Derive the "Continuity Equation" (more conservation) mass = Sg dV volnev desits S (1) change of mass with time: Surface 57 clearent Je Jedu -S svdŠ (2) mass flux through the surface : (1) and (2) must be equal!  $\frac{\partial}{\partial t} \int g dV = - \int g \vec{v} d\vec{s}$ 

Now use Gauss' theorem (a divergence theorem"):  

$$\int_{S} (\vec{q}) d\vec{S} = \int (\nabla \cdot \vec{q}) dV$$

$$\Rightarrow \frac{\partial}{\partial t} \int g dV = -\int \nabla \cdot (g \vec{v}) dV$$
Now pull  $\frac{\partial}{\partial t}$  into the integral:  

$$\Rightarrow \int \frac{\partial}{\partial t} g dV = -\int \nabla \cdot (g \vec{v}) dV$$
Drop integral over  $dV$ , because this small be  
true for any volume V:  

$$\Rightarrow \frac{\partial}{\partial t} S + \nabla \cdot (g \vec{v}) = 0$$
Continuity Equation (Eulerian form).

Using the product rule, me get  $\nabla \cdot \left( \vec{\nabla} \cdot \nabla \right) = \vec{\nabla} \cdot \left( \vec{\nabla} \cdot \vec{\nabla} \right) + g \left( \vec{\nabla} \cdot \vec{\nabla} \right)$ ... and back into the Continuity Eq:  $\Rightarrow \frac{\partial g}{\partial L} + (\vec{\nabla} \cdot \nabla) g = - g(\nabla \cdot \vec{\nabla})$ Define the co-moving (Lagrangian) derivative: carlesian.  $\nabla = \begin{pmatrix} \partial_{x} \\ \partial_{y} \\ \partial_{z} \\ \partial_{z} \end{pmatrix}$  $\frac{\mathcal{V}}{\mathcal{D}\mathcal{L}} = \frac{\partial}{\partial\mathcal{L}} + (\vec{v} \cdot \nabla)$ and use this in the above eq.:  $= \frac{Dg}{Df} = -g(\nabla \cdot \vec{\nabla}) \qquad \text{Continuity Eq.}$   $= \frac{Dg}{Df} = -g(\nabla \cdot \vec{\nabla}) \qquad \text{in Legrangian}$ form. Remark: for incompressible fluid: v·v=0 Conservation of momentum Basically follow the same approach as for mass,

but nows for momentum :  

$$\frac{\partial}{\partial t} \int_{V} (g\vec{v}) dV = -\int_{S} (g\vec{v}) \vec{v} d\vec{S} - \int_{S} Pd\vec{S}$$
Momentum flux pressure force  
Now can also nowile this as  

$$\frac{\partial}{\partial t} \int_{V} (g\vec{v}) dV = -\int_{S} (g\vec{v}\vec{v} + 1! \cdot P) d\vec{S}$$

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$$\int_{V} (g\vec{v}) dV = -\int_{V} \nabla \cdot (g\vec{v}\vec{v} + 1! \cdot P) dV$$

$$\implies \frac{\partial}{\partial t} (g\vec{v}) + \nabla \cdot (g\vec{v}\vec{v}) = -\nabla P$$
Now expand derivatives:  

$$S \frac{\partial\vec{v}}{\partial t} + \vec{v} \frac{\partial g}{\partial t} + \vec{v} \nabla \cdot (g\vec{v}) + S(\vec{v} \cdot q)\vec{v} = -\nabla P$$

$$= O (continuity eq.)$$

$$= \underbrace{\partial \vec{v}}_{\partial t} + (\vec{v} \cdot \nabla) \vec{v} = - \underbrace{\nabla P}_{S}$$
  
Euler equation (momentum conservation)  
in Eulerian form.

Using the Lagrangian derivative:  

$$\frac{D\vec{v}}{Dt} = -\frac{\nabla P}{S}$$
Monentran Equation  
in Lagrangian form.

Conservation of Energy  
Similar to mass and momentine eq., but for Sett 1  
which is the total energy density (nearns that  
Ctot is called specific (per mit mass) energy),  
Also need to add source terms as for the momentum  
equation (north from contraction / expansion, i.e. pressure):  

$$\frac{2}{2t}(getot) + \nabla \cdot [(getot + P)\nabla] = 0]$$
 Energy  
mith  $e_{tot} = e + \frac{\sqrt{2}}{2}$   
internal hintic (a specific energies)

 $\left\{\frac{D}{Dt}(e_{tot}) = -\frac{P}{S}(\nabla \cdot \vec{v}) - \frac{1}{S}\vec{v} \cdot \nabla P\right\}$ Energy Eq. in Lagrangian form.

laken together, me have 5 coupled PDEs, but 6 variables (1 mass, 3 moment, 1 energy) (8, v, P, et of) 

= System is not closed, and needs a closure eq. : the Equation of state (relates s, P, e) (EOS)