

# IDL

## Interactive Data Language

A Program Language used for data analysis & imaging processing  
(e.g., **astronomy**, atmospheric physics, medical imaging)



## Fun facts:

1. David Stern (Univ. Colorado; Lab for Atmospheric & Space Physics)  
1970s, PDP-12 -> SOL; He then found a company "research system Inc (RSI)"  
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Halley's Comet imaging processing

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License not free & expensive : (



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## Is it necessary to learn IDL ?

**IDL vs. IRAF, FORTRAN/C/C++, Python, Matlab**  
**other special packages (CIAO, STSDAS, AIPS)**



# Lots of Online Resources

## Get Started

IDL Commands and Syntax ([colorado.edu](http://colorado.edu))

A Slug's Guide to IDL (google pls)

A GUIDE TO IDL FOR ASTRONOMERS ([virginia.edu](http://virginia.edu))

Carl Heiles's Quick IDL Tutorial

## More advanced

Coyotes' Guide to IDL

David Fanning IDL programming Techniques

## Libraries

Astrolib (NASA)

JHU IDL Library

TeXtoIDL

## Many online personal collections:

Rob Dimeo's IDL programs

Robbie's IDL Programming

:

&

many many others



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**Have a clear logic**

**Have a good coding habit (always write instructions !)**



## From astrolib:

# GOOD

```

pro kstwo, data1, data2, D, prob
;+
; NAME:
;   KSTWO
; PURPOSE:
;   Return the two-sided Kolmogorov-Smirnov statistic
; EXPLANATION:
;   Returns the Kolmogorov-Smirnov statistic and associated probability
;   that two arrays of data values are drawn from the same distribution
;   Algorithm taken from procedure of the same name in "Numerical
;   Recipes" by Press et al., 2nd edition (1992), Chapter 14
;
; CALLING SEQUENCE:
;   kstwo, data1, data2, D, prob
;
; INPUT PARAMETERS:
;   data1 - vector of data values, at least 4 data values must be included
;           for the K-S statistic to be meaningful
;   data2 - second set of data values, does not need to have the same
;           number of elements as data1
;
; OUTPUT PARAMETERS:
;   D - floating scalar giving the Kolmogorov-Smirnov statistic. It
;       specifies the maximum deviation between the cumulative
;       distribution of the data and the supplied function
;   prob - floating scalar between 0 and 1 giving the significance level of
;          the K-S statistic. Small values of PROB show that the
;          cumulative distribution function of DATA1 is significantly
;          different from DATA2
;
; EXAMPLE:
;   Test whether two vectors created by the RANDOMN function likely came
;   from the same distribution
;
;   IDL> data1 = randomn(seed,40)      ;Create data vectors to be
;   IDL> data2 = randomn(seed,70)     ;compared
;   IDL> kstwo, data1, data2, D, prob & print,D,prob
;
; PROCEDURE CALLS
;   procedure PROB_KS - computes significance of K-S distribution
;
; REVISION HISTORY:
;   Written      W. Landsman          August, 1992
;   FP computation of N_eff      H. Ebeling/W. Landsman March 1996
;   Converted to IDL V5.0      W. Landsman September 1997
;   Fix for arrays containing equal values J. Ballet/W. Landsman Oct. 2001
;-
On_error, 2
if ( N_params() LT 4 ) then begin
  print,'Syntax - KSTWO, data1, data2, d, prob'
  return

```

## From myself:

# BAD

```

pro mapcube,ps=ps
device, decomposed=0
loadcolors
if keyword_set(ps) then setps_a4
if keyword_set(ps) then device,file='ha_on_hstv5.eps',/land,/col
;
;fits_read,'finalcube_v5.fits',cube,hdrf
fits_read,'finalcube_fcupdateflux.fits',cube,hdrf ;;; update flux, 10^-16 unit
sz = size(cube)
halpha = fltarr(sz[1],sz[2])
nii0=halpha
for i=0,sz[1]-1 do for j=0,sz[2]-1 do halpha[i,j] = total(cube[i,j,197:207]) ;;1
for i=0,sz[1]-1 do for j=0,sz[2]-1 do nii0[i,j] = total(cube[i,j,223:232]) ;;1.6
;
halpha = filter_image(halpha,fwhm=2.5,/all)
nii0 = filter_image(nii0,fwhm=2.5,/all)
;
;;; Make new header for the Halpha image
fits_read,'s100303_a023001_Hn3_100.fits',junk,asthdr
ctype1 = sxpar(asthdr,"CTYPE2")      ;; RA tan yori note, ctype1 of HST image is
ctype2 = sxpar(asthdr,"CTYPE3")      ;; DEC tan xori
mkhdr,newhdr,halpha,/im
sxaddpar,newhdr,"crval1",177.39710
sxaddpar,newhdr,"crval2",22.39601
sxaddpar,newhdr,"cdelt1",-2.7777778e-05 ;;;0.1"/pixel=> 0.1"/3600 = 2.77778e-5
sxaddpar,newhdr,"cdelt2",2.7777778e-05
sxaddpar,newhdr,"crpix1",29
sxaddpar,newhdr,"crpix2",39.5
sxaddpar,newhdr,"ctype1",CTYPE1
sxaddpar,newhdr,"ctype2",CTYPE2
sxaddpar,newhdr,"equinox",2000.0
;
writefits,'halpha_updateflux.fits',halpha,newhdr
hast0=halpha*100.
writefits,'halpha_updateflux_unit16std.fits',halpha*100.,newhdr
ok1=where(hast0 gt 1.)
hast=hast0-hast0 & hast[ok1]=hast0[ok1]
writefits,'halpha_updateflux_unit16std_gt1.fits',hast,newhdr
;
ok1=where(hast0 gt 0.8)
hast=hast0-hast0 & hast[ok1]=hast0[ok1]
writefits,'halpha_updateflux_unit16std_gt0.5.fits',hast,newhdr
halphatmp=hast/100.
;

```



## Commands (Procedures)

"IDL> procedure\_name, arg1, arg2, ..."

where arg1, arg2 can be either input or output.

## Function

"IDL> returnvalue = function\_name(arg1, arg2, ...)"

## Keywords

"IDL> procedure\_name, arg1, keyword=3"

/keyword is the same as keyword=1.



## Little things I wish someone told me before:

1. File name Case Sensitive (no capital letters!)  
e.g., case.pro, but not CASE.pro

> What this suggests to me  
> is that you cannot have a file that stores a routine and name it with  
> **capital letters**. That is

> dumb. Someone please tell me that I'm right???

Uh, well, I think this has been well known since  
about 1956. :-)

Cheers,

David

--

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Coyote's Guide to IDL Programming: <http://www.dfanning.com/>

Toll-Free IDL Book Orders: 1-888-461-0155



plotting examples: see demo

- ❖ see also **Online Resources** for fancier things you can to make prettier plots

- ❖ **IDL can handle 3D plots and animations quite well**

- ❖ Please read this to understand IDL color models,  
[http://www.geo.mtu.edu/geoschem/docs/IDL\\_Manuals/working\\_with\\_color.pdf](http://www.geo.mtu.edu/geoschem/docs/IDL_Manuals/working_with_color.pdf)  
device, decomposed=0 vs.  
device, decomposed=1



Read and write ASCII files: see demo

Little things I wish someone told me before:

2.

> When you open a file for writing, 80 columns is the default maximum  
> for historical reasons. (Are you old enough to remember  
> punch cards?)

e.g., in textopen.pro

```
openw, !TEXTUNIT, 'laser.tmp'
```

```
add
```

```
openw, !TEXTUNIT, 'laser.tmp', WIDTH=1200
```



Manipulating FITS files: see assignments

### Read FITS file:

```
"IDL> image = mrdfits('filename', 0, header, /fscale, /silent)"
```

(alternative: `fits_read,'filename',image,header,exten_no=0`)

idl will read the fits file into this idl name "image", can be any name.

" 'filename' " is the name of the fits file.

"0" means fits extension number. It's usually 0.

"header" will return FITS header info in idl

"/fscale" rescale to original data

"/silent" suppress info messages.

### Write FITS file:

```
"IDL> mwrfits, image, 'filename', header, /create"
```

"/create" overwrite existing file.

(alternative: `writefits`)

### Get FITS file information

```
"IDL> fits_info,'s100303_a023001_Hn3_100.fits'
```



**Start Demo**